



# MICROSCOPE: Firsts results of the Equivalence Principle test in space...before last release

PARIS 2019 – ACES Workshop



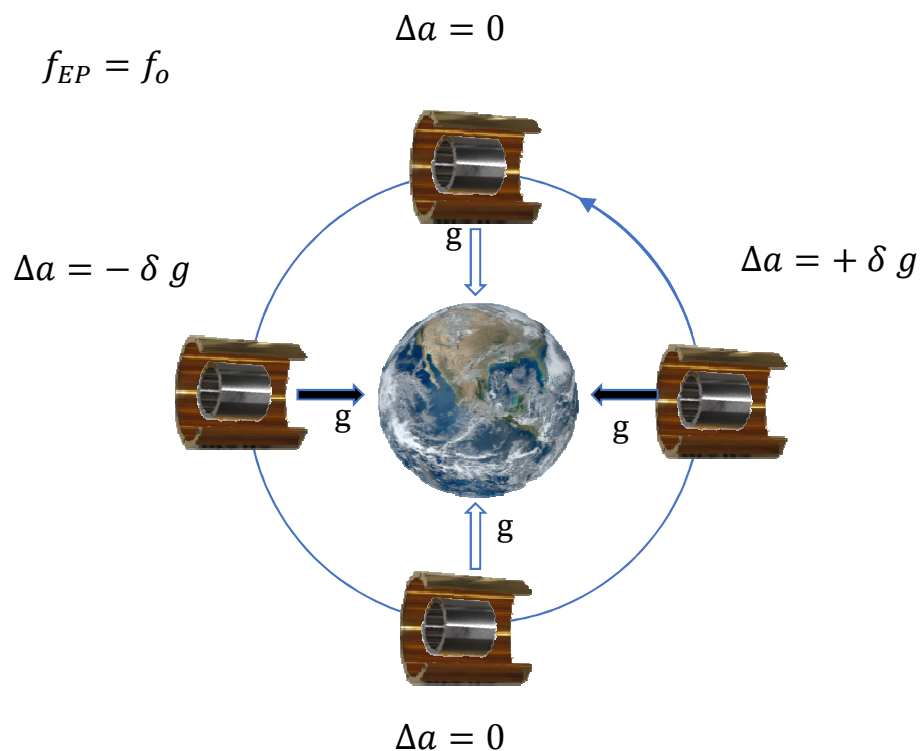
**Manuel RODRIGUES,**

**on behalf of the MICROSCOPE team**



# The « free-fall » test in space with MICROSCOPE resolution objective on WEP parameter: $\delta @ 10^{-15}$

$$\Delta a = a_1 - a_2 = \left( \frac{m_{g1}}{m_{i1}} - \frac{m_{g2}}{m_{i2}} \right) g$$



$m_g$  = gravitational mass



$m_i$  = inertial mass



Comparison of the acceleration of 2 bodies constrained to fall in the same gravity field:

$$\text{Eötvös Parameter } \delta = \frac{a_1 - a_2}{\frac{1}{2}(a_1 + a_2)} = \frac{\frac{m_{g1}}{m_{i1}} - \frac{m_{g2}}{m_{i2}}}{\frac{1}{2} \left( \frac{m_{g1}}{m_{i1}} + \frac{m_{g2}}{m_{i2}} \right)}$$

If  $\delta = 0$  :  $\Delta a = 0$

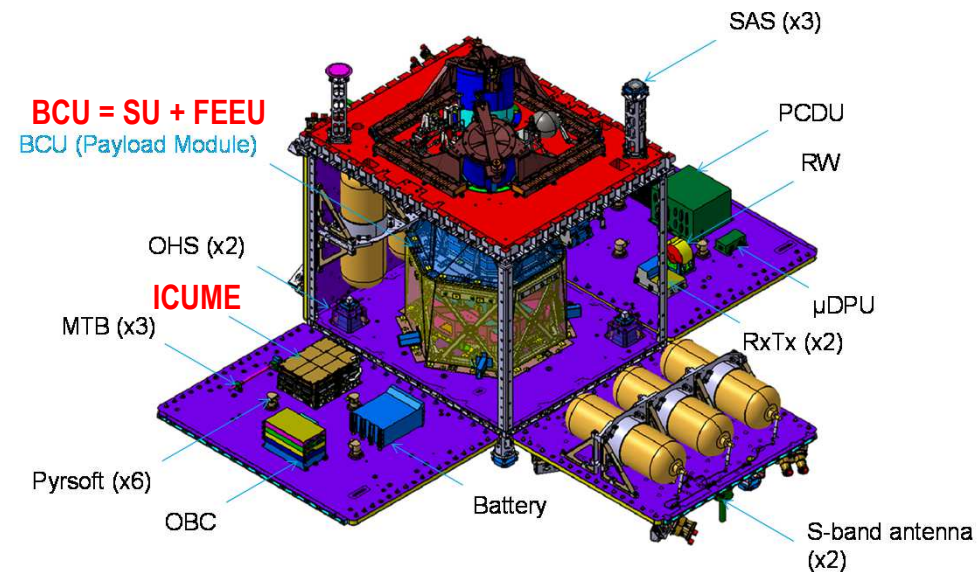
If  $\delta \neq 0$  :  $\Delta a \neq 0$  detection of a signal collinear to  $g$   
(same phase, same frequency)

$g(@710km) = 7.9m/s^2$

## The MICROSCOPE satellite

- ❖ Sun-synchronous orbit @ 710 km
- ❖ Several modes :
  - Inertial  $f_{EP} = \text{orbital frequency} = 1.7 \times 10^{-4} \text{ Hz}$
  - 2 rotation rates of S/C
    - $f_{EP} = 0.9 \times 10^{-3} \text{ Hz}$  &  $f_{EP} = 3.1 \times 10^{-3} \text{ Hz}$

- ❖ Cold Gas propulsion
- ❖ A space laboratory of 300kg
- ❖ 1,4 m x 1 m x 1,5 m
- ❖ Instrument in the BCU (Payload Thermal Cocoon Case) at the center of the satellite

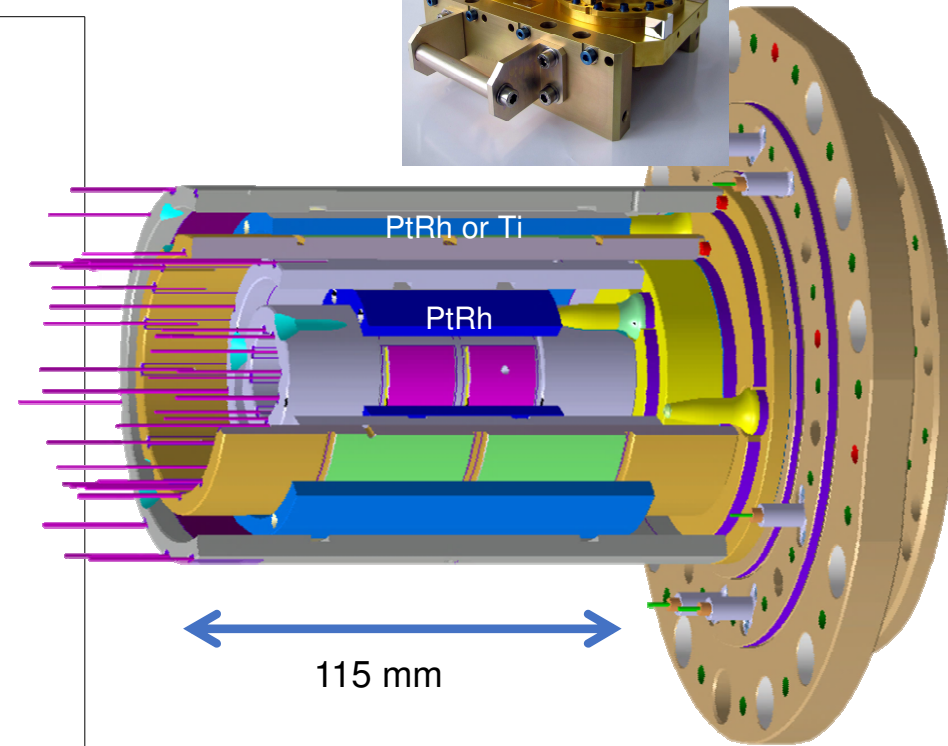


## Instrument : 2 double accelerometers for the test

2 Sensor Units on board which comprise each 2 concentric test-masses




SUEP : Sensor Unit with Ti / PtRh

SUREF : Sensor Unit with PtRh / PtRh, helps to get confidence on the overall performance and data process



# DRAG-FREE SATELLITE LABORATORY OF PHYSICS

## With capabilities of stimuli production:

-  ➤ linear or angular sine accelerations,
-  ➤ Test-masses displacements,
-  ➤ controlled thermal heaters (Off in science mode).

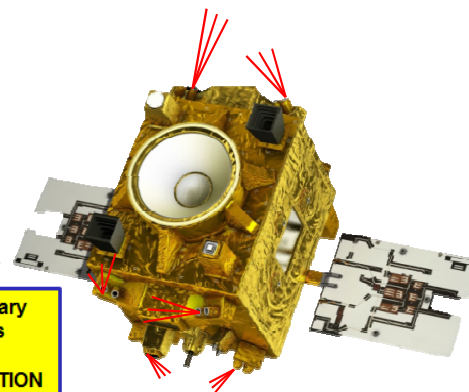
## Performance of drag-free

$$\Gamma(f_{EP}) < 3 \times 10^{-13} \text{ m/s}^2$$

$$\dot{\Omega}(f_{EP}) < 4 \times 10^{-12} \text{ rd/s}^2$$

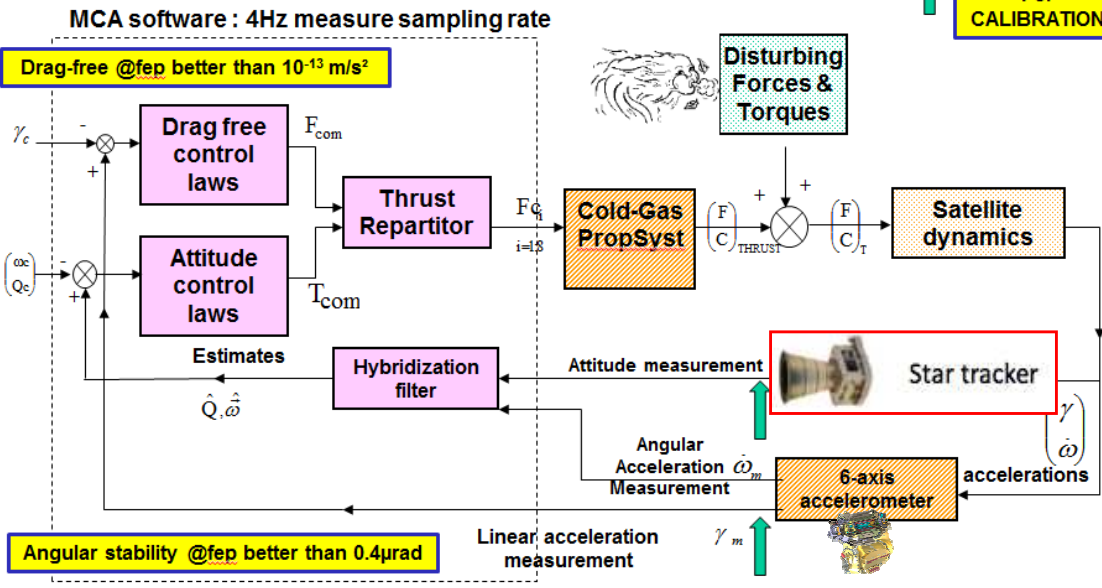
$$\Omega(f_{EP}) < 3 \times 10^{-10} \text{ rd/s}$$

$$\int \Omega < 1 \mu\text{rd}$$



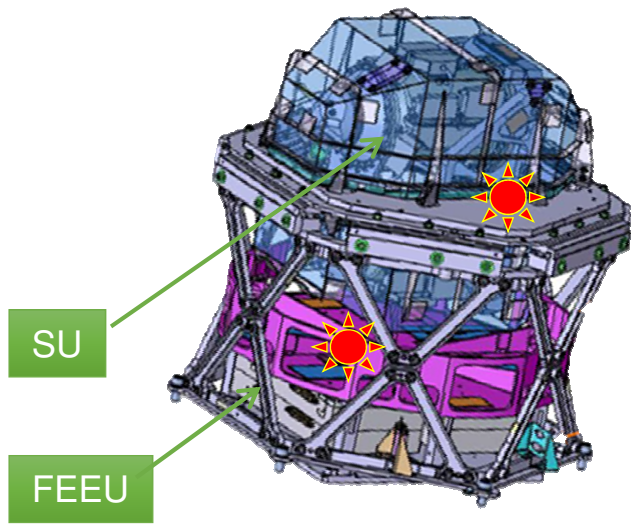
Secondary Inputs For CALIBRATION

Bandwidths: 12 SU control loops (1Hz) + 6 DFACS loop (0.1Hz) + 8 thruster loop (10Hz)

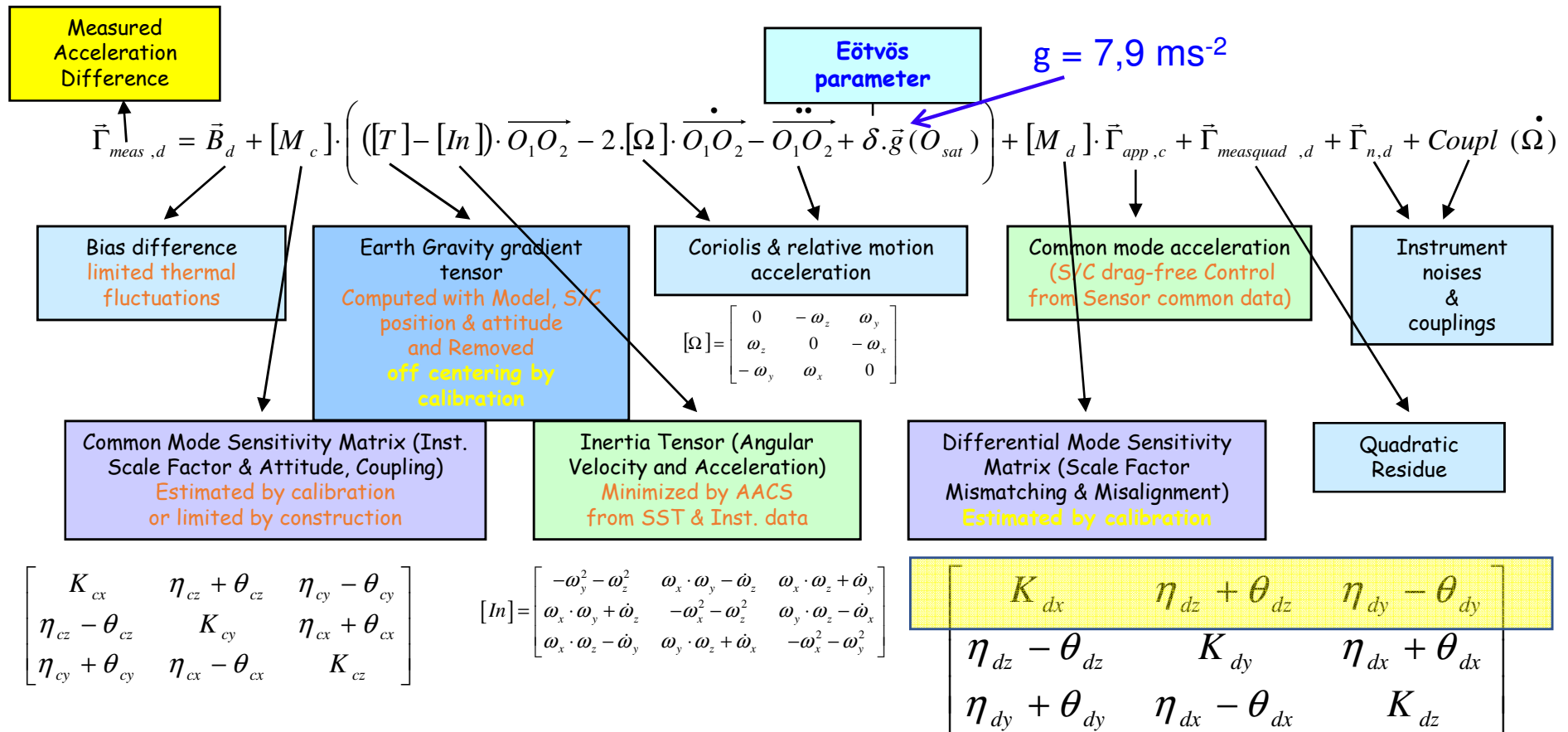


Drag-free @fep better than  $10^{-13} \text{ m/s}^2$

Angular stability @fep better than  $0.4 \mu\text{rad}$

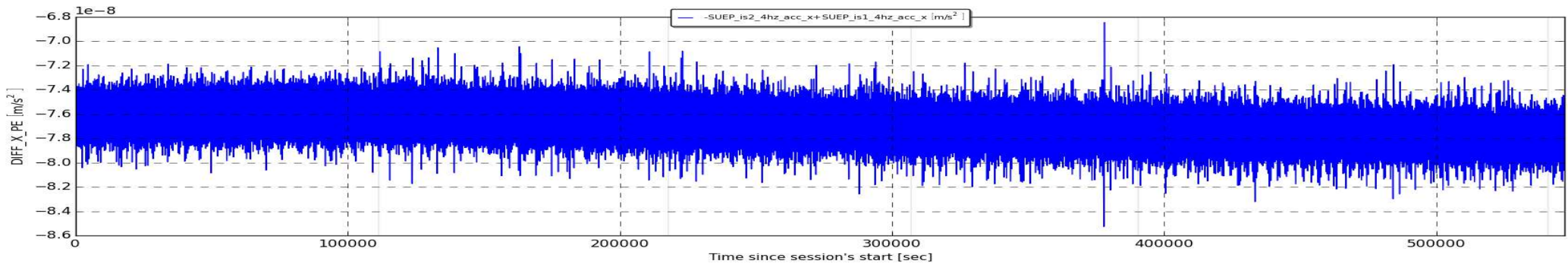


# The measurement: Earth, satellite, instrument and Physics contributions



**The measurement along the cylinder axis (X) = the main measurement**

## Measured time series and effect of in-flight calibration (session 234)

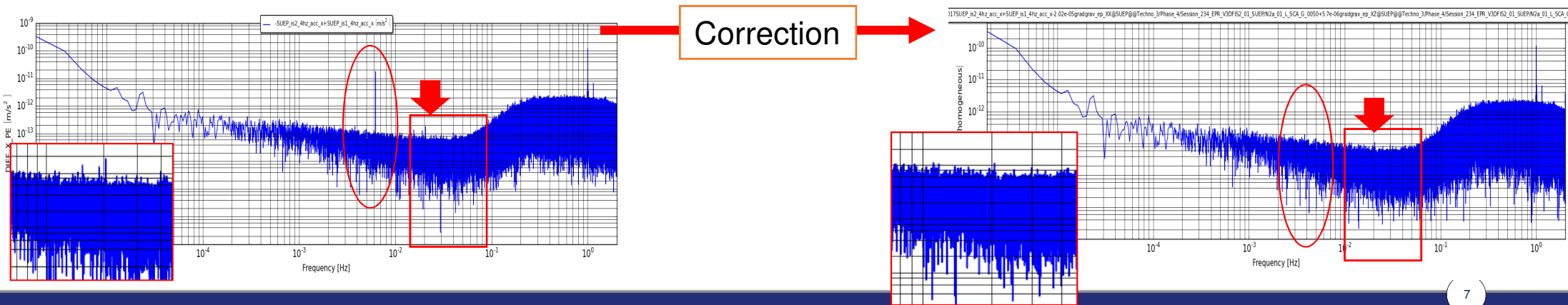


- ❖ Scale factor matching through a dedicated session before the EP session
- ❖ Test-mass off-centering estimated through the Earth's gravity effect at  $2f_{EP}$   
=> Correction of off-centering effects at all frequencies ( $f_{EP}$  and  $2f_{EP}$  included)

$$K_{dx} = 0.0085 \pm 6 \cdot 10^{-5}$$

$$\Delta X = (20.2 \pm 0.03) \mu\text{m}$$

$$\Delta Z = (-5.7 \pm 0.03) \mu\text{m}$$



# FIRST RESULTS PUBLISHED BASED ON 2 SESSIONS

SUEP

	dateDebut	nomFiche	Num Orb	contrainte Environnement	crit.	duree	etat	conso Gaz7n	conso Gaz7m	capacite Gaz7n	capacite Gaz7m
205	2017-02-13T13:59:55.846867		4321	NO_ECLIPSE_NO_LUNE	2	1.01295	E				
206	2017-02-13T15:40:18.833216	CAL_K1dxDFIS1_01_SUEP	4322	NO_ECLIPSE_NO_LUNE	2	5.07000	E				
207	2017-02-14T00:02:44.983178		4327	NO_ECLIPSE_NO_LUNE	2	1.01295	E				
208	2017-02-14T01:43:07.970959	CAL_K1dxDFIS2_01_SUEP	4328	NO_ECLIPSE_NO_LUNE	2	5.07000	E	2.9	3.3	6579.4	6614.3
209	2017-02-14T10:05:34.128091		4333	NO_ECLIPSE_NO_LUNE	2	3.07939	E	10	9.3	6369.1	6604.9
210	2017-02-14T15:10:44.141758	EPR_V3DFIS2_01_SUEP	4				E	176	151.3	6392.6	6453.3
211	2017-02-18T01:45:43.539435		4				E	5.4	4.2	6386.9	6448.7
212	2017-02-18T04:15:53.554441	EPR_V3DFIS2_01_SUEP	4				E	263.5	235.1	6123.1	6213.3
213	2017-02-23T09:55:00.000000		4464	NO_ECLIPSE_NO_LUNE	0	0.00000	E	0	0	6123.1	6213.3
214	2017-02-23T09:55:00.000000	TSNA	4464	NO_ECLIPSE_NO_LUNE	0	61.80639	E	0	3.3	6122.9	6209.7
215	2017-02-27T16:00:00.028541		4526	NO_ECLIPSE_NO_LUNE	2	1.01295	E	1.3	1.1	6121.7	6207.8
216	2017-02-27T17:40:23.014532	CAL_K1dxDFIS2_01_SUEP	4527	NO_ECLIPSE_NO_LUNE	2	5.07000	E	4.9	8.3	6116.9	6199.3
	7-02-28T02:02:49.160909		4532	NO_ECLIPSE_NO_LUNE	2	3.07939	F	10.4	11	6106.4	6187.8
218	7-02-28T07:07:59.169132	EPR_V3DFIS2_01_SUEP	4535	NO_ECLIPSE_NO_LUNE	2	120.00000	E	384.8	405.8	5721	5781.7
	7-03-08T13:19:57.511429		4655	NO_ECLIPSE_NO_LUNE	2	2.57703	E	3.7	4.9	5716.8	5776.4
220	2017-03-08T17:35:20.494387	CAL_tetadZDFIS2_01_SUEP	4658	NO_ECLIPSE_NO_LUNE	2	5.07000	E	3.9	7.9	5712.9	5768.3
							E	0.8	1.1	5712.1	5766.8
							E	3.6	8.3	5708.3	5758.3
							E	2.9	4	5705	5754
							E	11.9	15.7	5692.6	5738
							E	0.4	1.6	5691.9	5735.9
							E	3.3	7.3	5688.3	5728.3

$$\delta(T_i, P_t) = [-1 \pm 9(stat) \pm 9(syst)] \times 10^{-15}$$

From least square fit

218

Touboul et al. Phys. Rev. Letts. 119 231101 (2017)  
No evidence of violation >  $1,9 \times 10^{-14}$

- Over 120 orbits
- Statistical noise integrated over 120 orbits
  - Systematics evaluated with a majoring of SU temperature variations ( $15\mu K @ f_{EP}$ )
  - **90% of systematics come from upper bound limit on temperature variations**

SUREF

	18T14:22:59.978006		3944	NO_ECLIPSE_NO_LUNE	1	1.01295					
	18T16:03:22.968294	CAL_K1dxDFIS2_01_SUREF	3945	NO_ECLIPSE_NO_LUNE	1	5.07000					
	19T00:25:49.137973		3950	NO_ECLIPSE_NO_LUNE	1	3.07939					
174	2017-01-19T05:30:59.159261	EPR_V2DFIS2_01_SUREF	3953	NO_ECLIPSE_NO_LUNE	1	120.00000	E	81.1	67.5	6720	6750.3
	2017-01-27T11:42:57.925815		4073	NO_ECLIPSE_NO_LUNE	1	1.51531	E	1	0.6	6719	6749.6
176	2017-01-27T14:13:07.942964	EPR_V2DFIS2_01_SUREF	4074	NO_ECLIPSE_NO_LUNE	1	82.00000	E	56	48.4	6662.9	6701
	2017-02-02T05:39:19.100109		4156	NO_ECLIPSE_NO_LUNE	1	2.57703	E	1.8	2	6661	6699
178	2017-02-02T09:54:42.094912	CAL_tetadZDFIS2_01_SUREF	4159	NO_ECLIPSE_NO_LUNE	1	5.07000	E	3.1	2.8	6657.8	6696.2
										6695.5	6692.4
										6688.3	6674.5
										6673.8	6668.5
184	2017-02-03T16:36:41.576425	CAL_K21xx_02_SUREF	4178	NO_ECLIPSE_NO_LUNE	1	10.00000	E	5.1	5.3	6631.4	6668.5

$$\delta(P_t, P_t) = [+4 \pm 4(stat) \pm 8(syst)] \times 10^{-15}$$

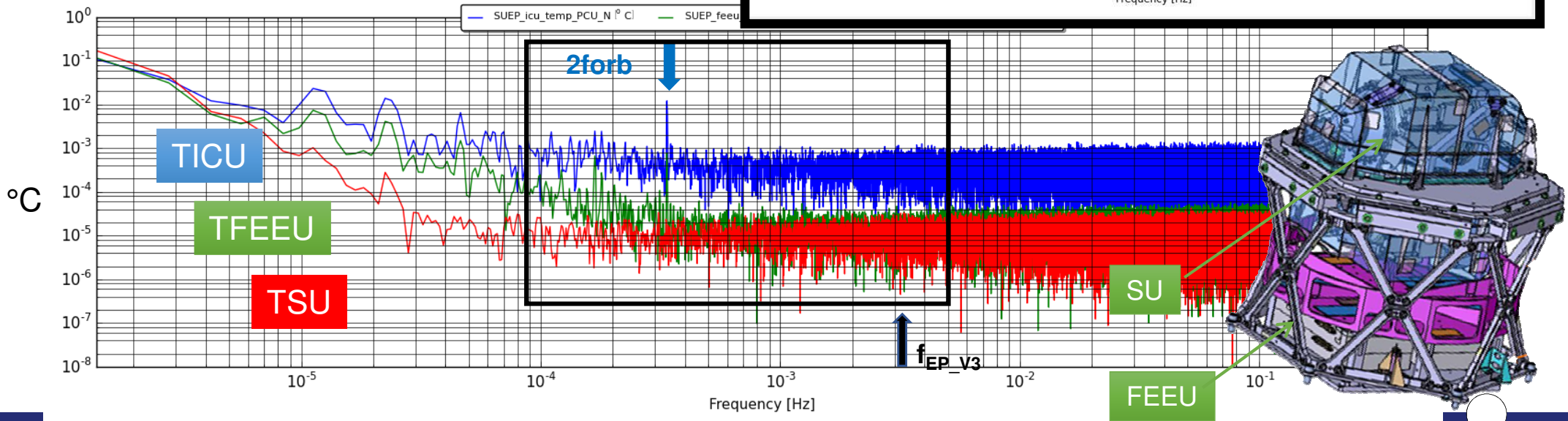
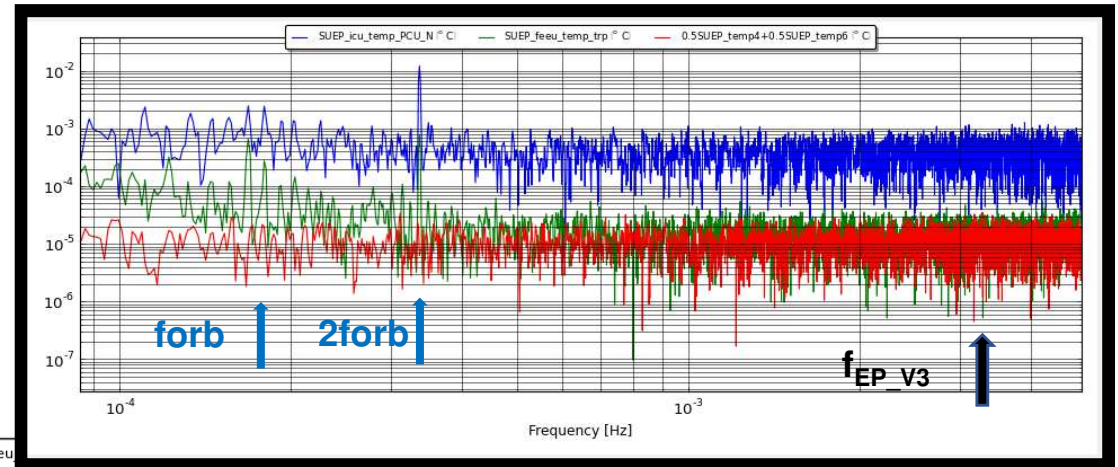
Touboul et al. Classical and Quantum Gravity, Vol. 36, N. 22, Oct 2019  
<https://dx.doi.org/10.1088/1361-6382/ab4707>

- Over 62 orbits
- Statistical noise integrated
  - Systematics evaluated with temperature measurements and evaluation of sensitivity



# Session 218 – SUEP EPRV3- Typical FFT of the temperatures (SU, FEEU, ICU)

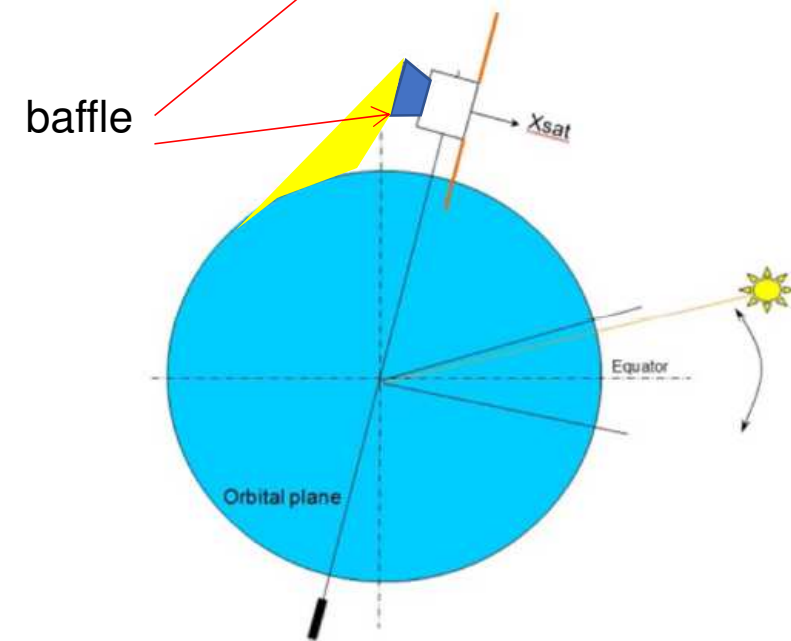
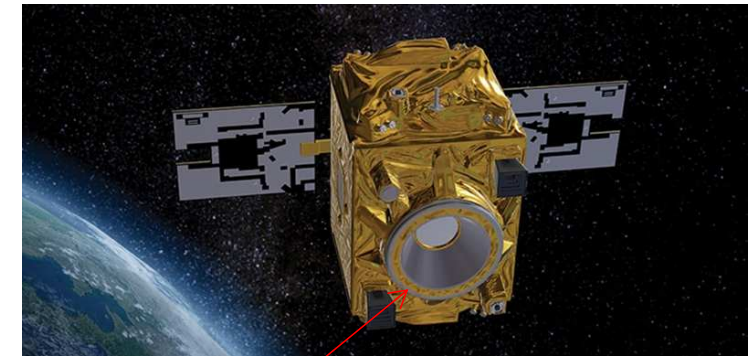
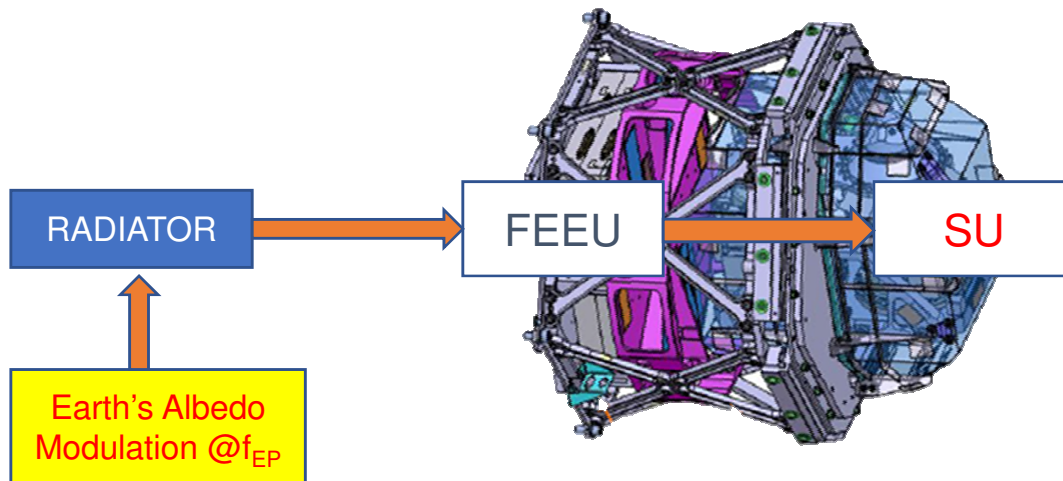
- ❖ The disturbances are mainly at  $f_{orb}$  &  $2f_{orb}$  in the FEEU. No signal at  $f_{EP}$
- ❖ No peak at  $f_{orb}$  in the ICU (only  $2f_{orb}$ ): the ICU is not in the BCU
- ❖ No significant temperature signal in the SU at all for  $f > f_{orb}$



## THERMAL TESTS IN FLIGHT

### OBJECTIVE:

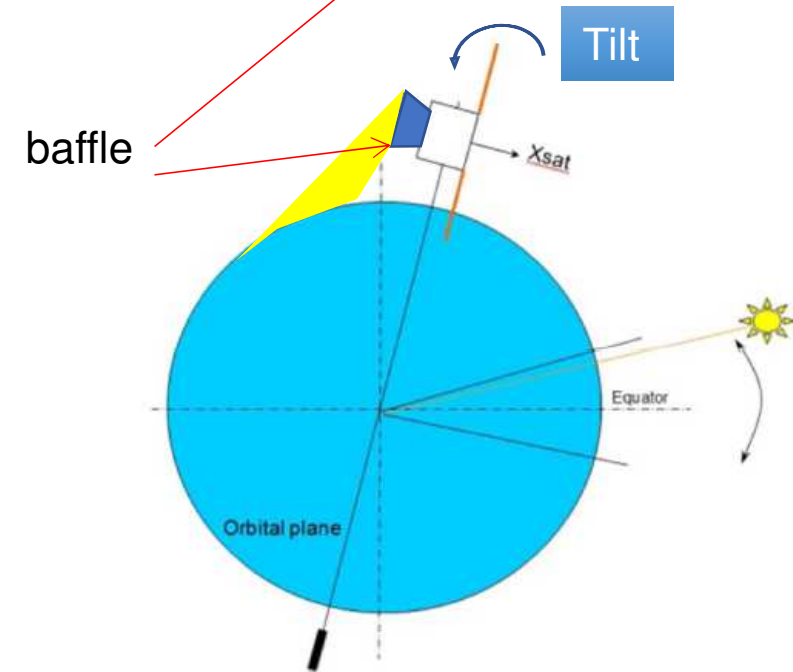
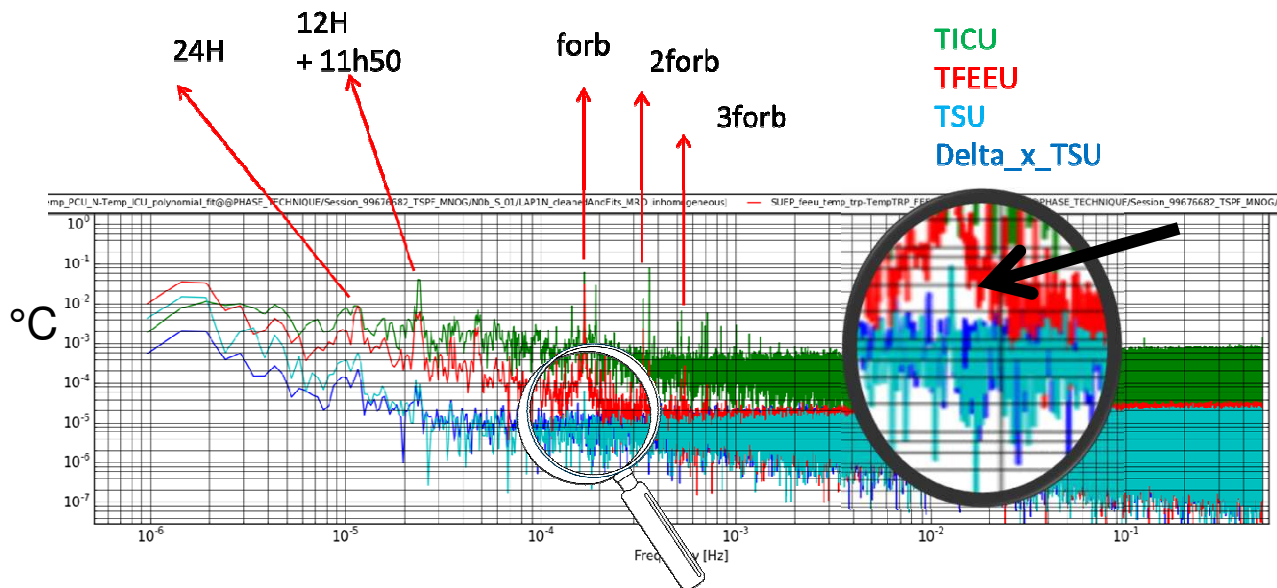
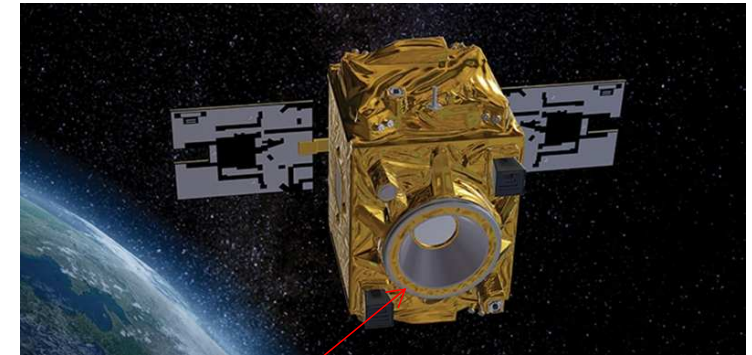
- To confirm the thermal model : temperature variations in the sensor (SU) is dominated by thermal conduction coming from the radiator through the FEEU plate
- To evaluate the thermal filtering from Radiator to FEEU and from FEEU to SU



# THERMAL FILTERING CHARACTERISATION

A specific session (*SPICHO*) of 460 continuous orbits was performed to amplify the Earth's thermal flux at orbital frequency by tilting the satellite

The result is awesome :  $\left( \frac{\Delta T_{FEEU}}{\Delta T_{SU}} \right)_{1.7 \times 10^{-4} \text{ Hz}} = 500$



## Thermal sensitivity & systematics

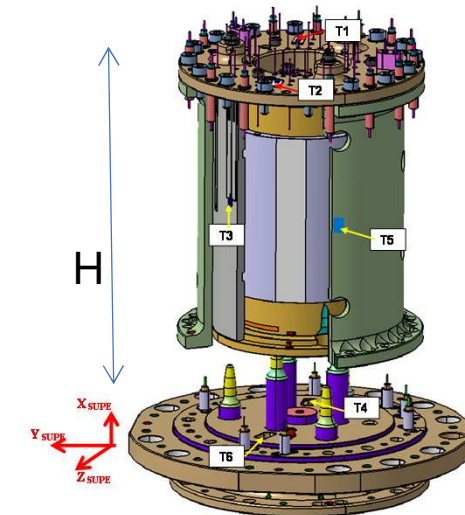
	SUREF	SUEP
Sensitivity to $T_{SU}$ at $f_{EP}$ [ $\text{ms}^{-2}\text{K}^{-1}$ ]	$3.9 \times 10^{-9}$	$4.3 \times 10^{-9}$
Sensitivity to $T_{FEEU}$ , at $f_{EP}$ [ $\text{ms}^{-2}\text{K}^{-1}$ ]	$5 \times 10^{-11}$	$7 \times 10^{-11}$

$$\Delta\Gamma_{dx}(syst_{therm}) = \left| \frac{\partial\Gamma_{dx}}{\partial\nabla T_{SU}} \Delta\nabla T_{SU} \right| \cdot H + \left| \frac{\partial\Gamma_{dx}}{\partial T_{SU}} \Delta T_{SU} \right| + \left| \frac{\partial\Gamma_{dx}}{\partial T_{FEEU}} \Delta T_{FEEU} \right|$$

$$|\Delta\nabla T_{SU}| \cdot H < |\Delta T_{SU}| \ll |\Delta T_{FEEU}| \text{ in all sessions}$$

In all science sessions :  $|\Delta T_{SU}|$  is dominated by the probe noise but thanks to The thermal characterization :

$$\Delta T_{SU} < 0.1 \mu\text{K} \text{ (instead of } 15 \mu\text{K of the upper limit in PRL)}$$



$$\nabla T_{SU} \cdot H = \frac{1}{2} (T1 + T2 - T6 - T4)$$

$$T_{SU} = \frac{1}{2} (T6 + T4)$$

# Some first resulting physics from 2017/2019 papers

## MICROSCOPE and modified gravity: generic 5<sup>th</sup> force model

JB, P. Brax, G. Métris, M. Pernot-Borràs, P. Touboul, J.-P. Uzan, 2018, PRL 120 141101

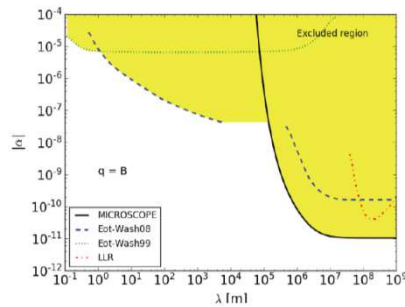
Yukawa potential

$$V_{ij}(r) = -\frac{Gm_i m_j}{r} \left(1 + \alpha_{ij} e^{-r/\lambda}\right)$$

$$\alpha_{ij} = \alpha \left(\frac{q}{\mu}\right)_i \left(\frac{q}{\mu}\right)_j$$

WEP violation

$$\eta = \alpha \left[ \left(\frac{q}{\mu}\right)_{Pt} - \left(\frac{q}{\mu}\right)_{Ti} \right] \left(\frac{q}{\mu}\right)_E \left(1 + \frac{r}{\lambda}\right) e^{-\frac{r}{\lambda}}$$



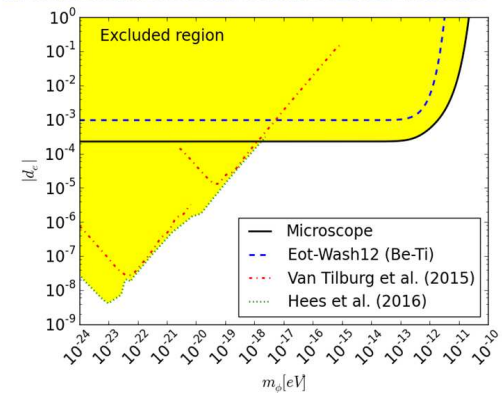
## Massive dilaton coupled to EM only

JB, P. Brax, G. Métris, M. Pernot-Borràs, P. Touboul, J.-P. Uzan, 2018, PRL 120 141101

$$d_g = d_{m_i} = 0$$

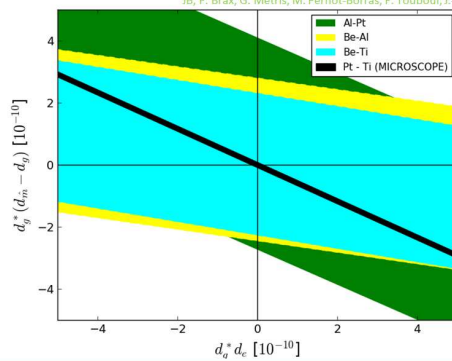
$$D_e \propto (d_e)^2$$

Van Tilburg, Hees: oscillations of the fine structure constant in a spectroscopic analysis of two isotopes of dysprosium



## Light dilaton vs MICROSCOPE

JB, P. Brax, G. Métris, M. Pernot-Borràs, P. Touboul, J.-P. Uzan, 2018, PRL 120 141101



New limits on Dilaton & Yukawa

MICROSCOPE Mission: First Constraints on the Violation of the Weak Equivalence Principle by a Light Scalar Dilaton  
Bergé et al. PRL 120, 141101

## Some first resulting physics from 2017/2019 papers

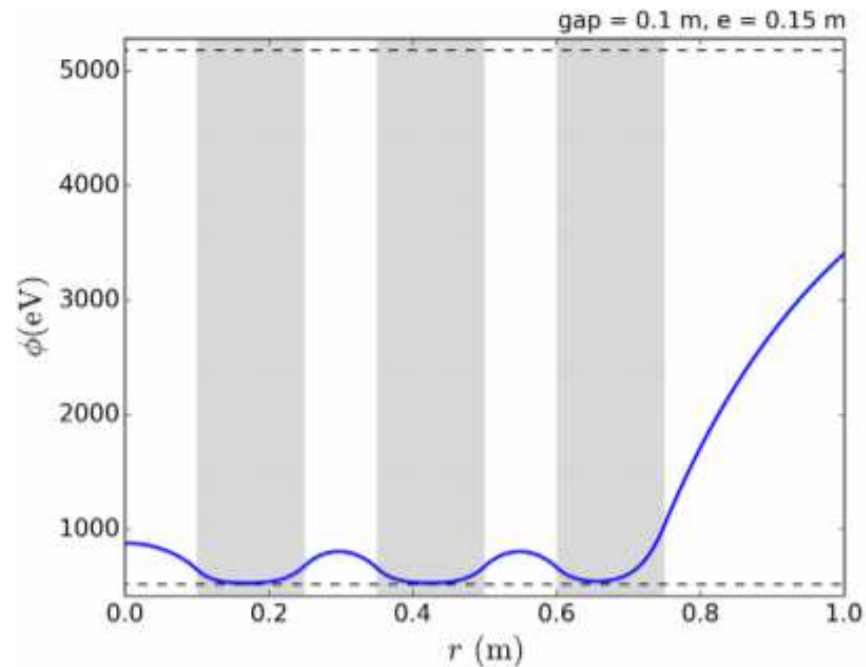
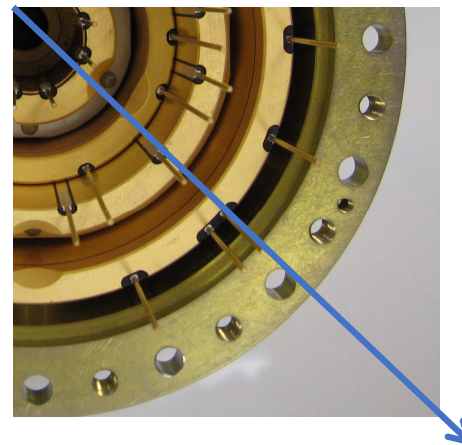


FIG. 20. Radial profile for three nested cylinders of thickness  $e$  with the same matter density. Cylinders are delimited by the shaded regions and separated by a distance  $gap$ . The  $\phi_{\min}$  values are represented by the horizontal segments.

**MICROSCOPE & Chameleonic forces:  
Screening effects of local mater modelling**

General study of chameleon fifth force in  
gravity space experiments  
Pernot-Borràs et al. PRD 100, 084006 (2019)



## Some first resulting physics from 2017/2019 papers

MICROSCOPE limits on the strength of a new force with comparisons to gravity and electromagnetism

Pierre Fayet, Phys. Rev. D **99**, 055043 –  
Published 28 March 2019

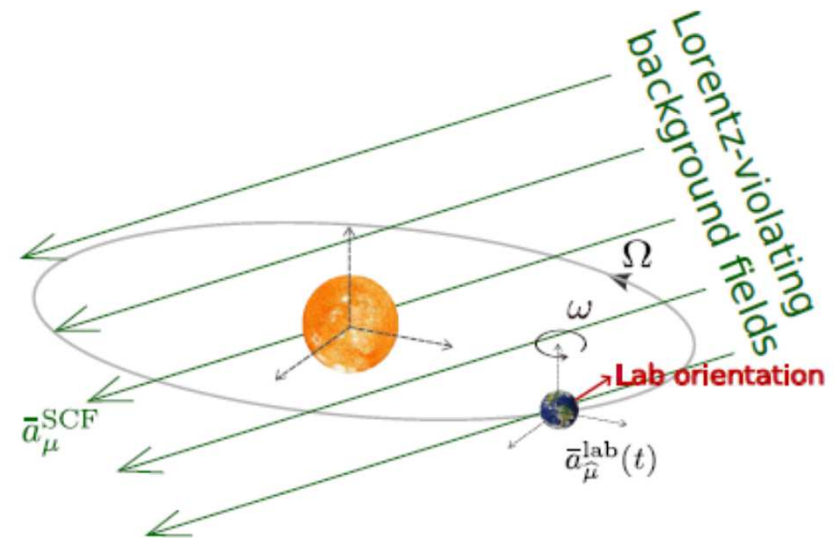
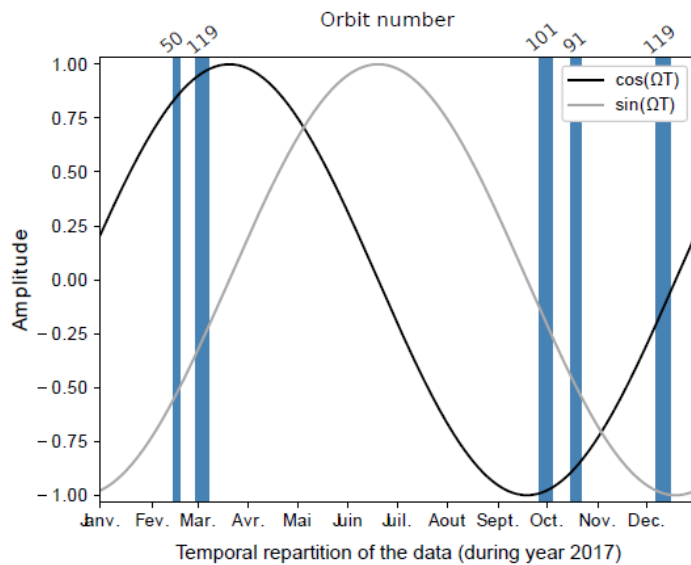
Improved limits on the strength of a new long-range force by one order of magnitude

MICROSCOPE limits for new long-range forces and implications for unified theories  
Pierre Fayet, Phys. Rev. D **97**, 055039 –  
Published 26 March 2018

## Some first resulting physics from 2017/2019 papers

New test of Lorentz invariance using the MICROSCOPE space mission  
Phys. Rev. Lett.  
Pihan-le Bars et al, October 2019 accepted

5 sessions analyzed  
New constraints on SME coefficients  
SEE NEXT PRESENTATION





## SUMMARY OF MICROSCOPE FLIGHT

Launched on the 25<sup>th</sup> of April 2016, switched off on the 16<sup>th</sup> of October 2018 : 2.5 year

20% (2768 orbits) dedicated to science

5% of thermal tests

19% of technological experiment

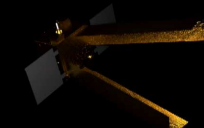
*1 orbit = 6000 sec*

10 papers under preparation with the final consolidated result: submission by end of 2019

⇒ Release by mid 2020

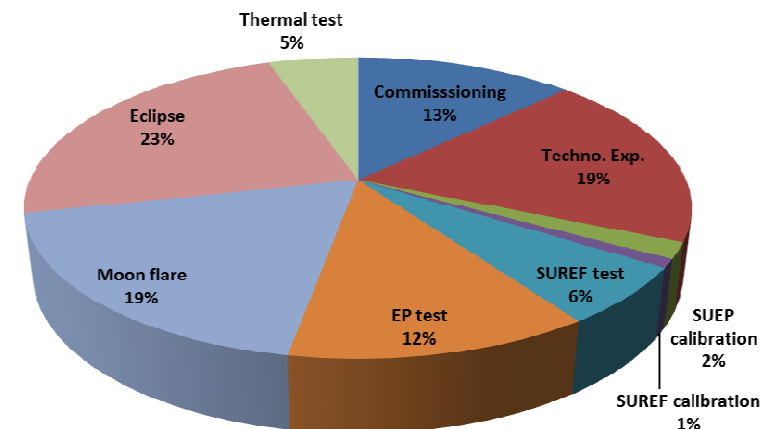
⇒ Data opened to public after release

© ONERA: Adaptative optics – OCA telescope



ONERA  
THE FRENCH AEROSPACE LAB

MICROSCOPE : 13 193 orbits (2016-2018)



OCTOBER 17th – SWICHT OFF SAT FEST



THANK YOU FOR YOUR ATTENTION

